## SO/PHI data request form

(Cruise phase + first science orbit; SO/PHI-Team internal version)

## CLV of the facular contrast using SO/PHI in combination with SDO/HMI

N. Krivova, K.L. Yeo, S.K. Solanki, F. Kahil

MPS

## Science case (stay on one slide):

Please also state, why is PHI needed; why is the science unique?

- Uncertainties in the facular contrast, in particular its dependence on B and  $\mu$ , limit the reliability of the spectral irradiance models. This has consequences for our understanding of solar influence on climate.
- Current issues with the available facular contrasts(B,  $\mu$ ):
  - ➤ Observations: B is usually not available; observations close to the limb are uncertain;
  - ➤ Models: mostly 1D (poor closer to the limb) with no direct link to B; 3D models start appearing (observational validation needed).
- By observing the same regions on the Sun from different viewing angles, SO/PHI and SDO/HMI can provide B directly without foreshortening, which offers a unique opportunity to constrain the facular contrast close to the solar limb.

## Requirements/data

Besides best guess requirements, you may also list minimum requirements on the data

- Type of solar feature: faculae/network/QS
- HRT or FDT: FDT; if possible also HRT at a few orbital positions, especially when further away from the Sun
- Physical parameters needed (available: B\_LOS, vector B, v\_LOS, I\_c, raw data): B\_LOS, I\_c
- Total length of observation: 5-10 minutes daily or every few days throughout the whole window
- Cadence (maximum 1 dataset/min): 1 dataset/min
- Pointing needs (disc centre, limb, active region location, particular μ): For HRT: SW+PW around disc center, NW westward of disc centre
- Orbit needs (spatial resolution/co-rotation/angle to Earth/angle to other spacecraft): SW + PW more important; if available, data from the cruise phase would also be helpful
- Total number of datasets:  $\sim 1 \text{ set/min } \times (5-10 \text{ min/day}) \times (10-30 \text{ days}) \approx 50 300 \text{ sets}$
- Full frame 2k x 2k or partial frame 1kx1k, 0.5kx0.5: full frame
- Full resolution or 2x2, 4x4 binned data: full
- noise level (default 10<sup>-3</sup>): default
- Co-observations with other instruments: SDO/HMI (should be available)
- Special requests: For FDT: see pointing needs