



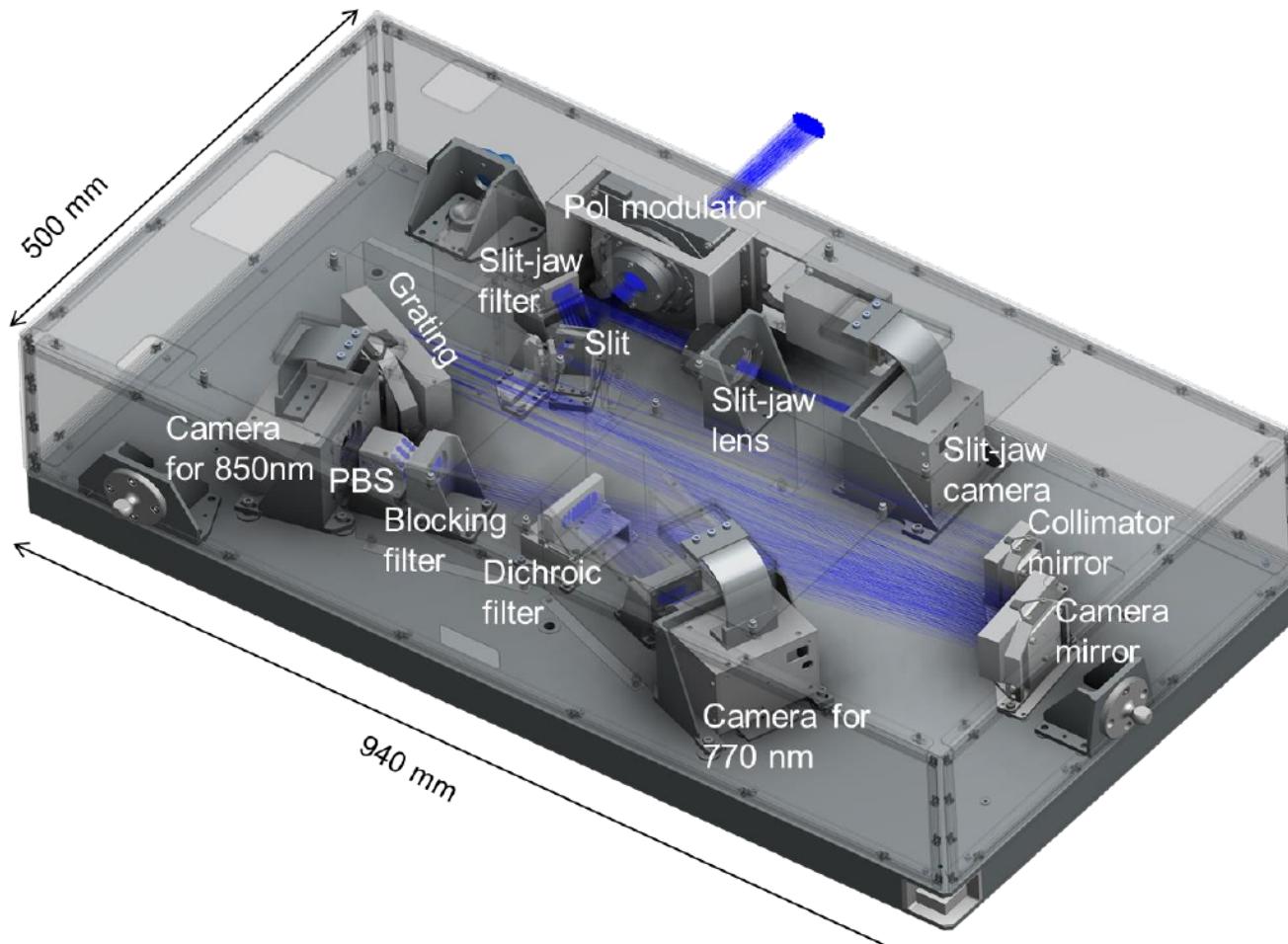
Sunrise Chromospheric Infrared spectropolarimeter (SCIP)

Basic specifications



	requirement	specification
wave length	To seamlessly cover the photosphere and the chromosphere simultaneously	1 . Ca II line, 850nm band (846.6-854.9 nm) 2. K I line, 770nm band (765.5-771.6nm) (Simultaneous observation of the 2 bands)
spatial resolution	To resolve dynamic MHD events in the chromosphere	0.2" (diffraction limit@854nm, = Hinode/SOT's imaging resolution), 0.094"/pixel
cadence		Stokes IQUV mode: 1-10 s per slit position Stokes I only: 45s / Full FOV
spectral resolution	To observe magnetic field in the chromosphere	$\lambda/\Delta\lambda = 2 \times 10^5$
polarimetric accuracy		requirement: 0.05% of Ic ($B_{LOS}=10G$ @Ca II 854nm) objective: 0.03% of Ic ($B_{LOS}=5G$ @Ca II 854nm)
FOV	To cover dynamical activity (> supergranular size)	58" (slit) x 58" (scan)

Optical layout



Properties of spectral lines

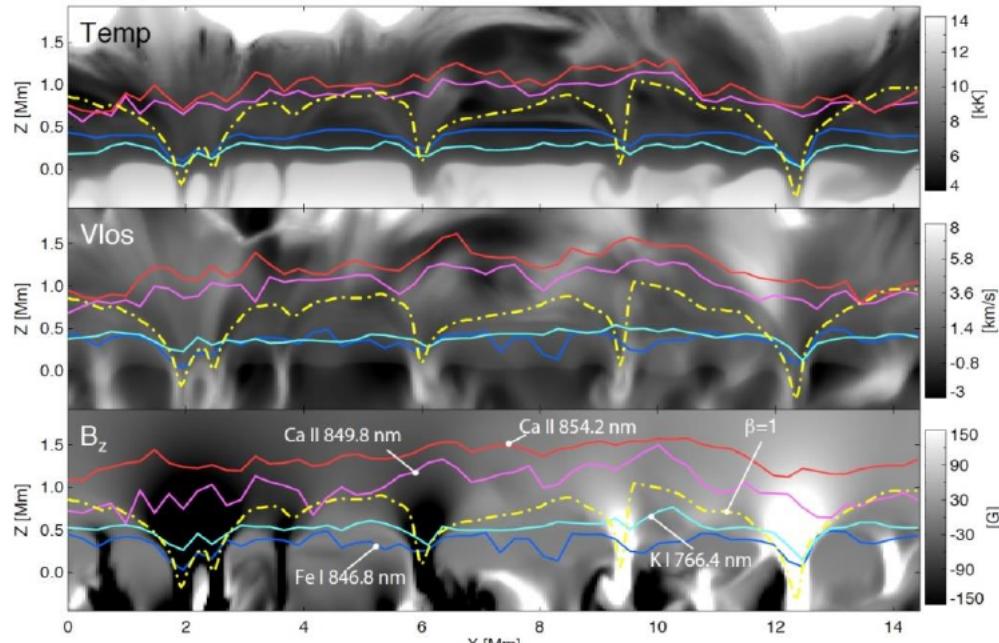


Properties of the main spectral lines in SCIP observation

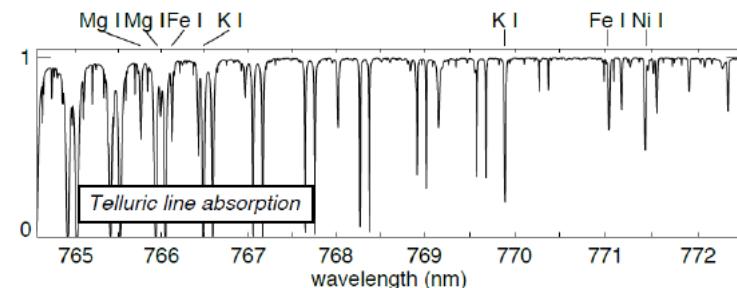
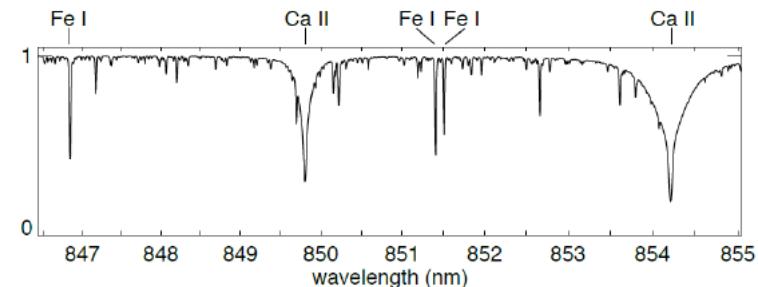
ph: photosphere, ch: chromosphere

line	Fe I	Ca II	Fe I	Fe I	Ca II	Mg I	Mg I	Fe I	K I	K I	Fe I	Fe I
λ (nm)	846.84	849.80	851.41	851.51	854.21	765.76	765.92	766.12	766.49	769.90	771.04	771.43
g_{eff}	2.50	1.07	1.83	0.75	1.10	1.25	1.75	1.60	1.16	1.33	1.50	1.50
layer	ph	ch	ph	ph	ch	ph	ph	ph	upper ph	upper ph	ph	ph

Properties of spectral lines



Line formation heights in SCIP observation predicted by 3D MHD simulation & NLTE line synthesis.



Spectral atlas in SCIP observing wavelength range.

Ca II, 850 nm band

K I, 770 nm band

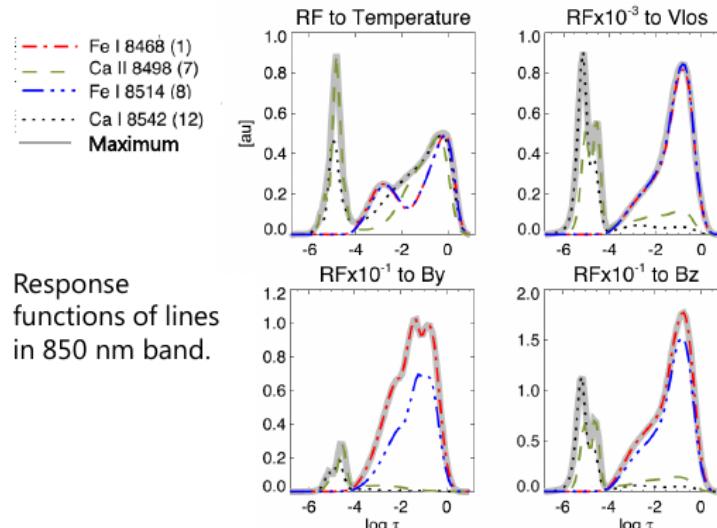
Properties of spectral lines



- **Ca II line, 850 nm band**

(Quintero Noda+2016,2017)

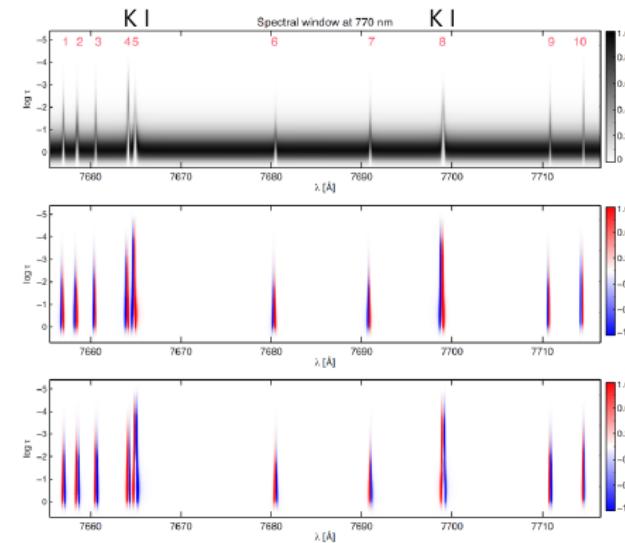
- Ca II lines :
sensitive to T & V_{LOS} from the
photosphere to the middle
chromosphere
- Fe I lines:
increase sensitivity at the
photosphere



- **K I line, 770 nm band**

(Quintero Noda+2017, 2018)

- K I lines :
sensitive to V_{LOS} & B
around the upper photosphere
→ complement 850nm band
- Telluric line contamination in K I D₂
becomes weak in stratosphere



Observing modes

SP		
	Normal mode (Stokes IQUV)	Rapid mode (Stokes I only)
FoV along the slit	58" (0.0936"/pix)	
FoV in the scanning direction	N_{step} steps (N=1 to 620, 1 step = 0.0936")	
Wavelength range (nm, in air)	CH1	846.6 - 854.7
	CH2	765.5 - 771.6
Integration time (*) total integration time = 512 ms x $N_{\text{pmu}} \times N_{\text{slit}}$	N_{pmu} cycles ($N_{\text{pmu}}=2$ to 20, 1 cycle = 0.512 ms) N_{slit} repeats ($N_{\text{slit}}=1$ to 100)	10 msec
Scan duration (sec, including overhead)	$(0.512 \times N_{\text{pmu}} \times N_{\text{slit}} + 0.128) \times N_{\text{step}}$	$0.064 \times N_{\text{step}}$
SJ		
FoV	60" x 60" (0.0936"/pix)	
Filter pass-band (in air)	770.5 ± 0.5 nm	
Exposure time	10 msec (TBD)	
Cadence	64 msec	

Black: fixed

Red: selectable/adjustable

Polarization sensitivity

(1σ sensitivity)

Integration time	Spatial summing	SP-CH1 (850 nm)	SP-CH2 (770 nm)
1.024 sec	no sum	1.4×10^{-3}	1.0×10^{-3}
	2 pix sum	1.0×10^{-3}	7×10^{-4}
5.12 sec	no sum	6×10^{-4}	4×10^{-4}
	2 pix sum	4×10^{-4}	3×10^{-4}
10.24 sec	no sum	4×10^{-4}	3×10^{-4}
	2 pix sum	3×10^{-4}	2×10^{-4}

Note:

- Disk-center observation
- Spatial sampling: $0.0936'' \times 0.0936''$ (no sum), $0.0936'' \times 0.187''$ (2 pix sum)
- Wavelength sampling: $\lambda/\Delta\lambda = 2 \times 10^5$
- Photon budget (single beam)
 - CH1: 1.1×10^6 e⁻/pix·sec
 - CH2: 2.3×10^6 e⁻/pix·sec
- Dual-beam combined
- Polarization efficiency: 0.5 for QUV