

Compression Analysis

Preliminary Status Report

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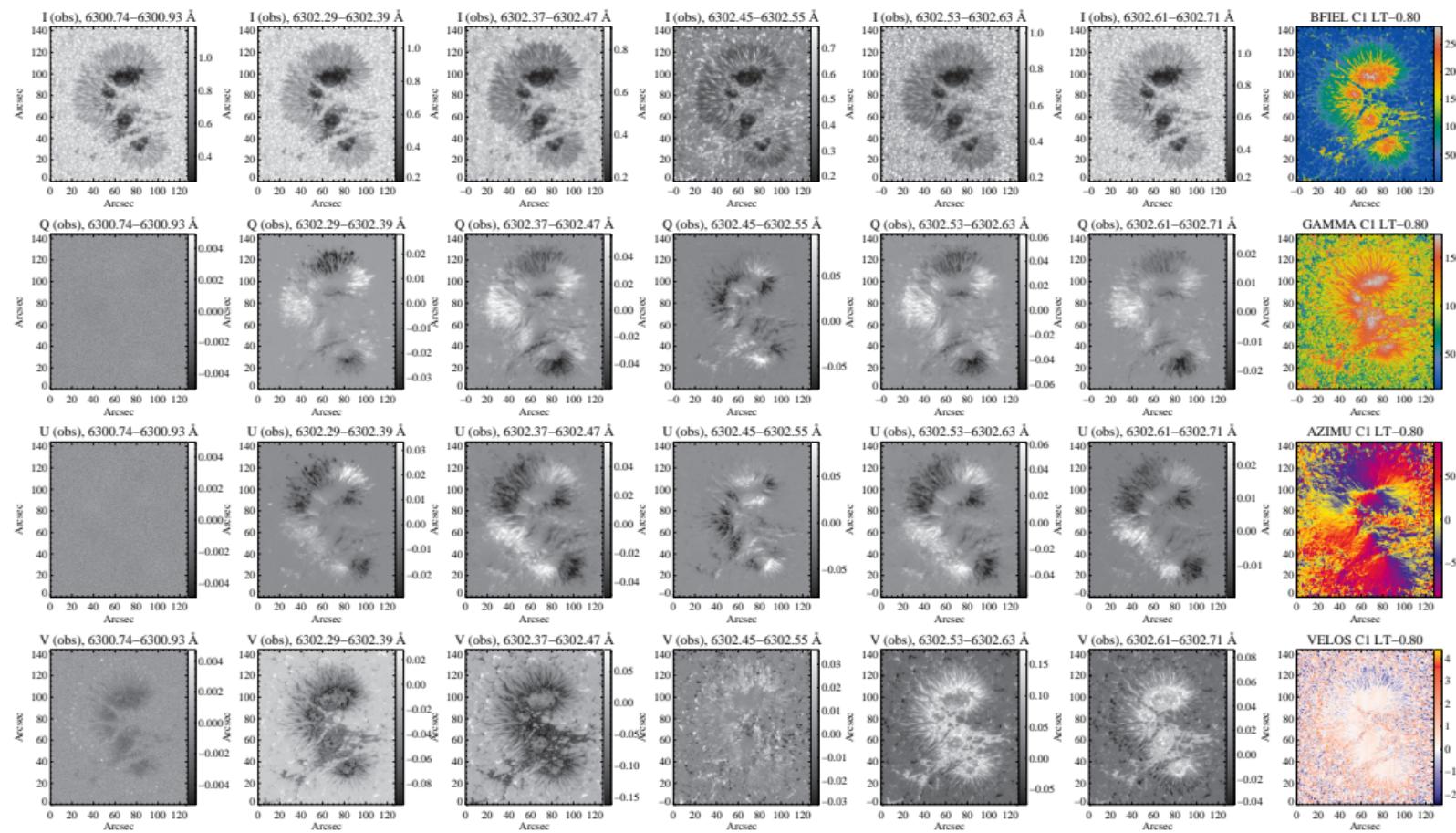
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ESAC, Jul 7-10 2014



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Milne-Eddington “Compression”

ME-inversions: Pros

- Raw-data combined to 24 Stokes images (24 bit)
- on-board inversion: reduced to 1 – 5 images, 8 – 10 bit
 - extremely efficient LOSSY compression: factor 10 – 50
- almost no loss of science (ideal case only)
 - avoids errors in inversion caused by compression artefacts

ME-inversions: Cons

- very risky:
 - requires perfect on-board data reduction (dark images, flat fielding, fringe removal, polarimetric calibration, cross-talk removal)
- very complex: Sunrise II IMaX experience

Most sophisticated realization ever for on-board data volume reduction!

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Lossy raw-data / Stokes-Vector Compression

- (1) Raw data Compr.
- (2) Stokes Vector Comp.

- allows to perform more sophisticated data analysis on ground
(e.g. coupled inversions, height dependent atmospheres)
- allows better control on data reduction
(fringe removal, PSF deconvolution, polarimetric calibration)
- Lossy compression noise must be below data noise
(goal: 10^{-3})
- Compression ratios much lower than ME-compression

(3) Lossy compression of ME results

(3) Compression of ME parameter maps

- acceptable compression noise level higher than for Stokes data (0.5%, TBD)
 - Higher compression factors possible
- Influence on helioseismology data: to be investigated
- Influence on field extrapolations: to be investigated
- Achievable compression factors: TBD (science driven)
 - SOPHISM simulations required
- Azimuth Ambiguity: lossy compression may lead to wrong results (to be investigated)

Stokes Vector Compression

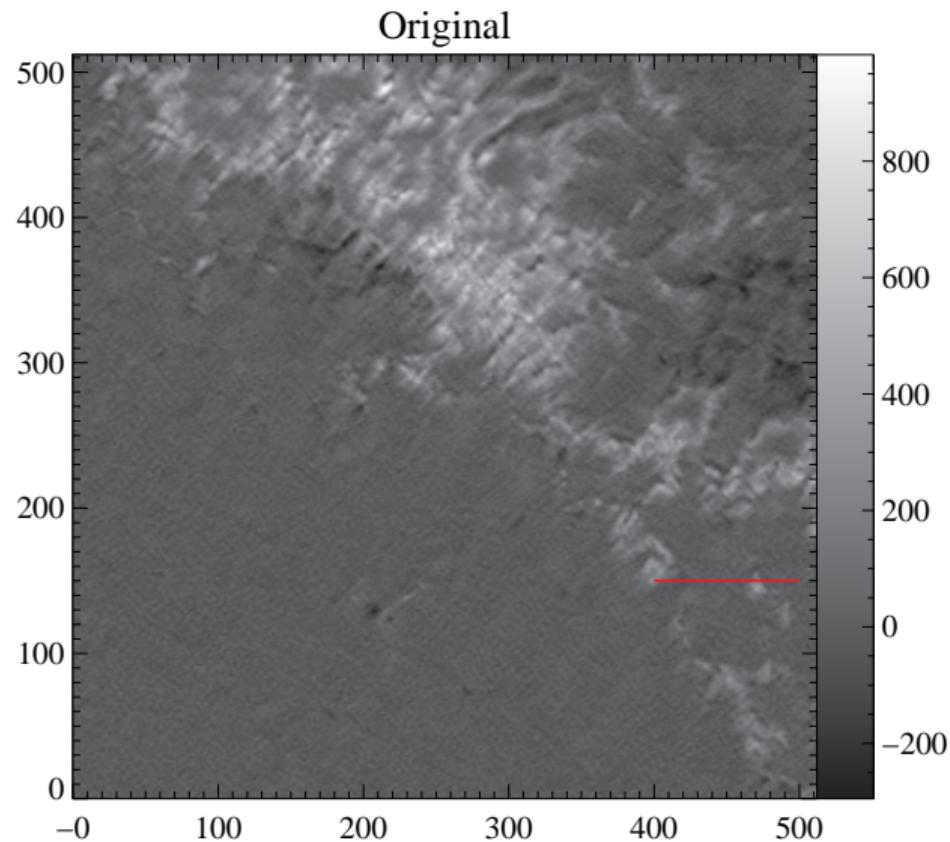
Procedure

- ① Use IMaX L2, Stokes V image
(+40mÅ from line core)
 - ② reduce BPP
 - ③ compress image with given ratio
 - ④ decompress image
- determine file size ratio
- compute difference: map, RMS

Compression Code

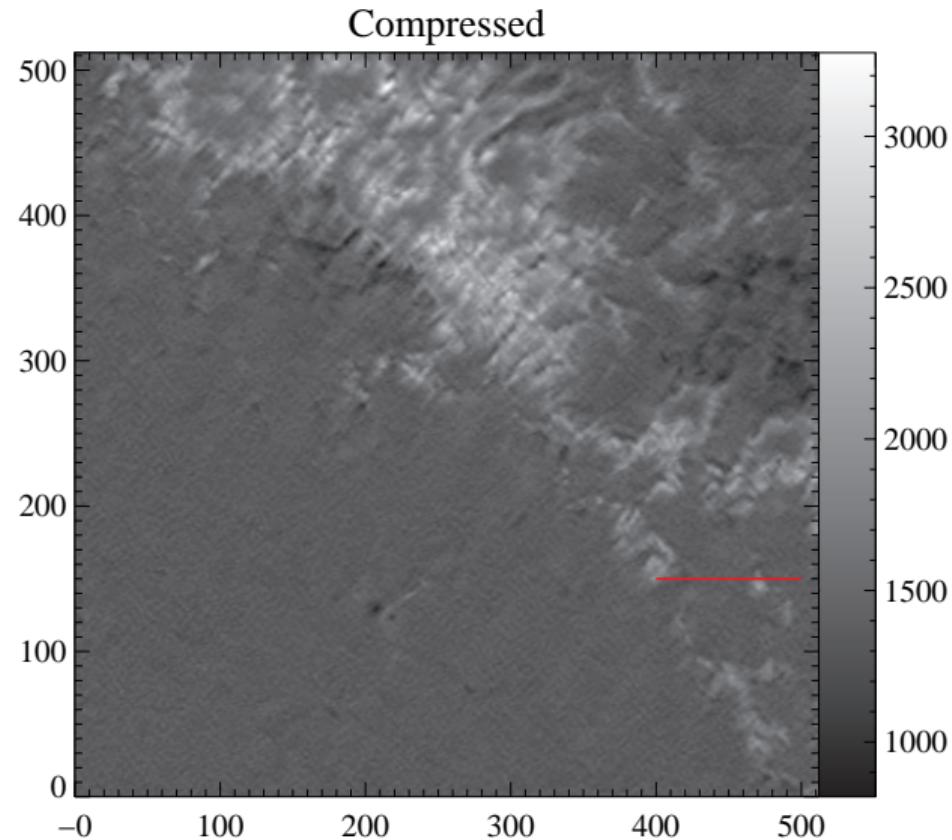
- provided by IDA Braunschweig
- several lossy algorithms available
- same code will run on-board (FPGA)
- implemented to SOPHISM (June 2014)

IMaX Stokes V
Original

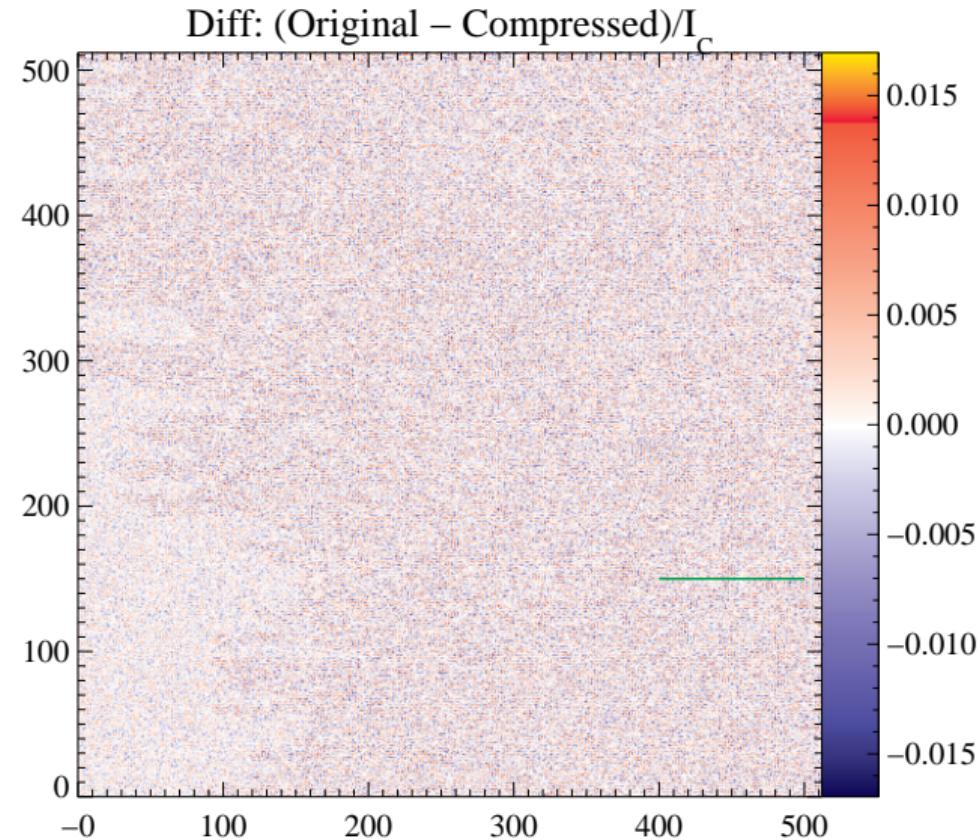


IMaX Stokes V

Compressed:
12 BPP, ratio: 8.0
size reduction
 $\times 10.67$,
RMS=3.02E-03

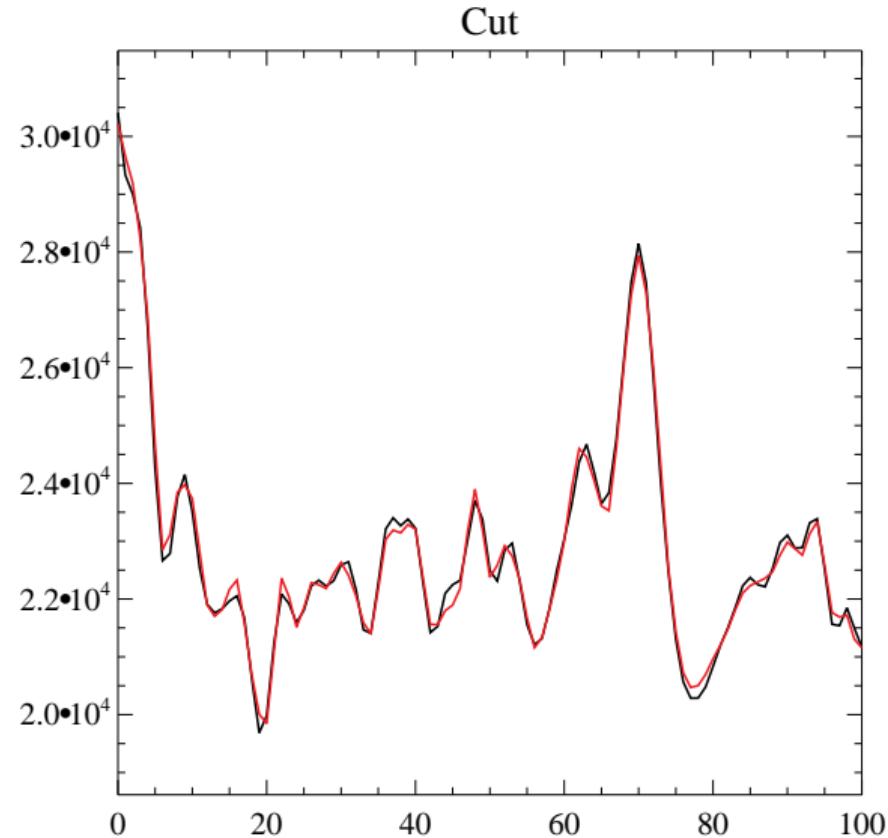


IMaX Stokes V
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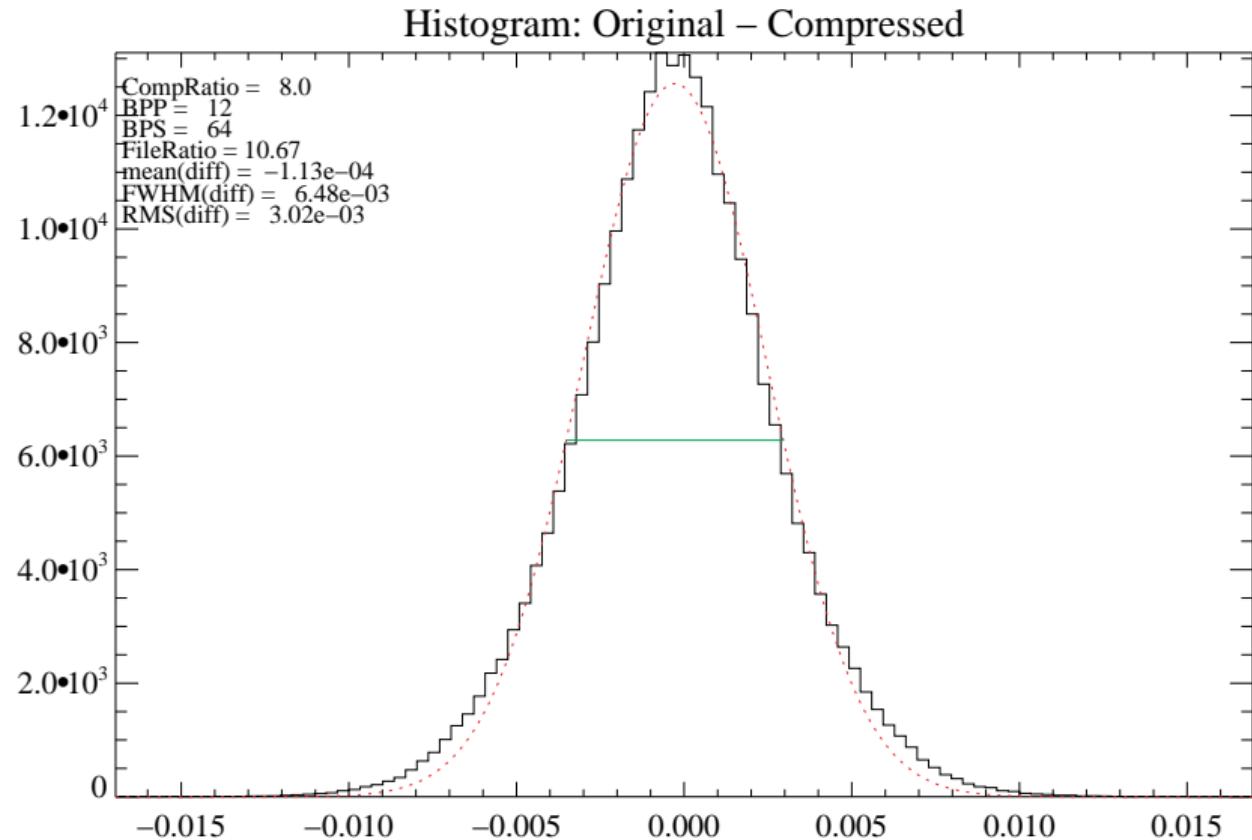


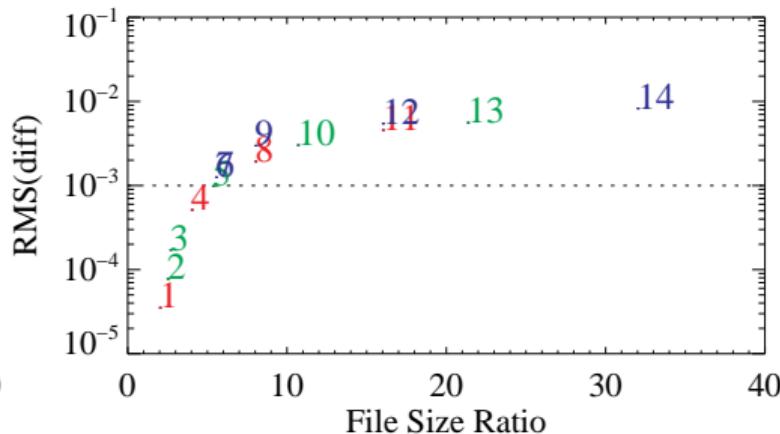
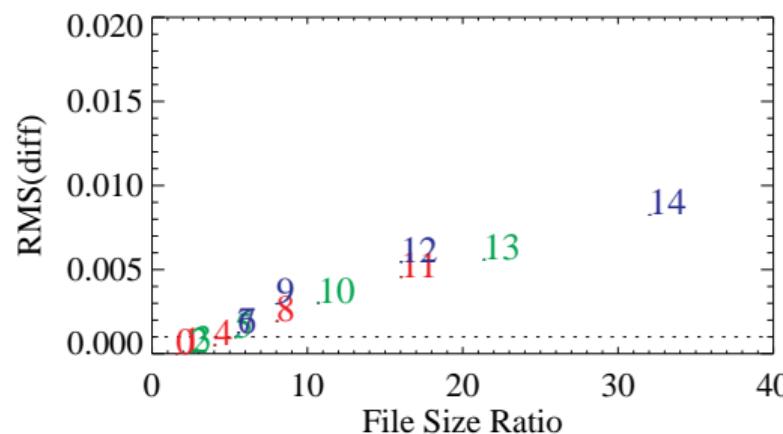
IMaX Stokes V

Cut: Original,
Compressed
12 BPP, ratio: 8.0
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IMaX Stokes V
Original -
Compressed
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SUMMARY: StokesV, IMaXL2

FileRatio	CompRatio	BPP	BPS	mean(diff)	FWHM(diff)	RMS(diff)
0	1.54	1.0	16	64	0.00e+00	0.00e+00
1	2.00	2.0	16	64	-2.91e-06	8.96e-05
2	2.49	1.0	12	64	-1.49e-08	2.23e-04
3	2.67	2.0	12	64	-5.48e-06	2.77e-04
4	4.00	4.0	16	64	-6.97e-06	9.51e-04
5	5.33	4.0	12	64	-1.32e-04	2.09e-03
6	5.53	1.0	8	64	3.24e-08	3.56e-03
7	5.53	2.0	8	64	3.24e-08	3.56e-03
8	8.00	8.0	16	64	-2.11e-06	4.43e-03
9	8.00	4.0	8	64	-9.41e-05	5.56e-03
10	10.67	8.0	12	64	-1.13e-04	6.48e-03
11	16.00	16.0	16	64	-6.90e-07	9.36e-03
12	16.00	8.0	8	64	2.93e-04	1.19e-02
13	21.33	16.0	12	64	-2.52e-05	1.23e-02
14	32.00	16.0	8	64	7.92e-04	1.86e-02

Summary - Stokes Vector Compression

Lossless Compression

Compression factors: $\times 1.4 - \times 1.9$

BPP Compression

noise level for bit truncation always higher than lossy compression at same file size ratio

Lossy Compression

achievable compression ratios for compression noise levels of $\leq 10^{-3}$ (Stokes *QUV*, all raw-data images)

→ $\approx \times 4 - \times 5$

for compression noise levels of 0.5% (maybe sufficient for Stokes *I*)

→ $\approx \times 10$

Preliminary Results! SOPHISM analysis still to be done!

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Further Compression Tests

Critical Issues:

- helioseismology / extrapolation aspect of compression
- bad pixel treatment
- vector magnetic field maps:
Compression without disambiguation possible?

New Test Setup:

- use SOPHISM
- test full pipeline
- **only then results have the necessary level of realism**

Test Matrix

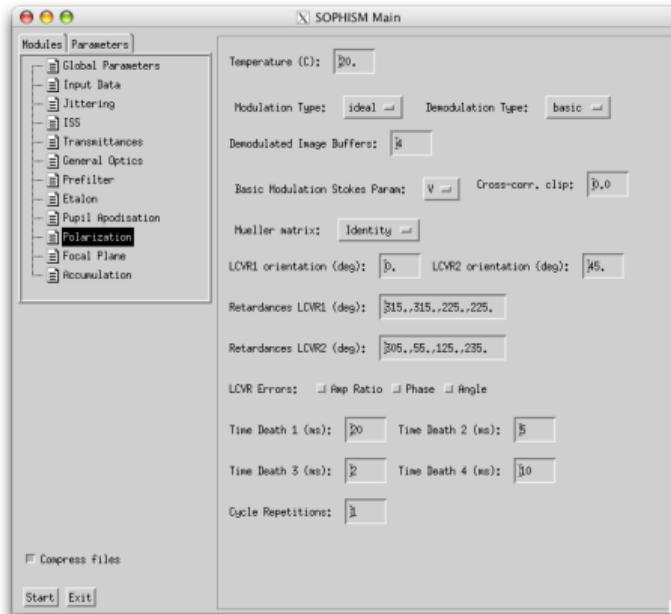
- raw data, Stokes vectors, ME-maps
- FDT, HRT
- lossless, $\times 2, 4, 8, 12$
- bit depth variation: 16, 14, 12, 10
- compression algorithms
- data with 2 different noise levels

Total: ≈ 800 tests

Time Table:

- computations: end of September
- results: end of October

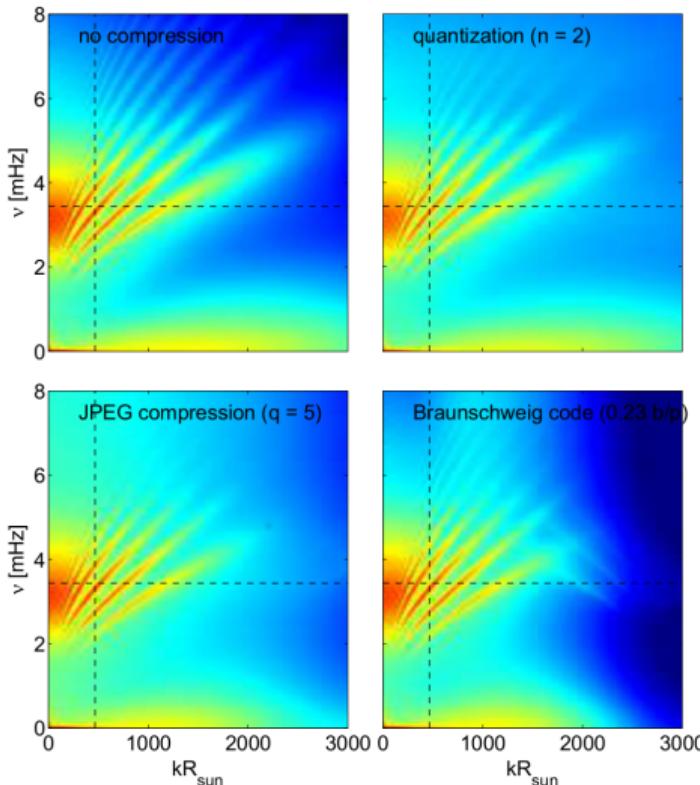
Solo PHI Instrument Simulator (SOPHISM)



(Julian Blanco, Univ. Valencia)

- Software IDL-coded SO/PHI simulator
- Main effects simulated in quasi-independent modules:
 - Jittering+ISS
 - Polarization modulation
 - Spectral profile
 - Optical aberrations
 - Pupil apodization
 - FPA, Accumulation
 - Modulation scheme
 - Solar evolution
 - Compression
 -
- Modules generate result files

Helioseismic Power Spectra

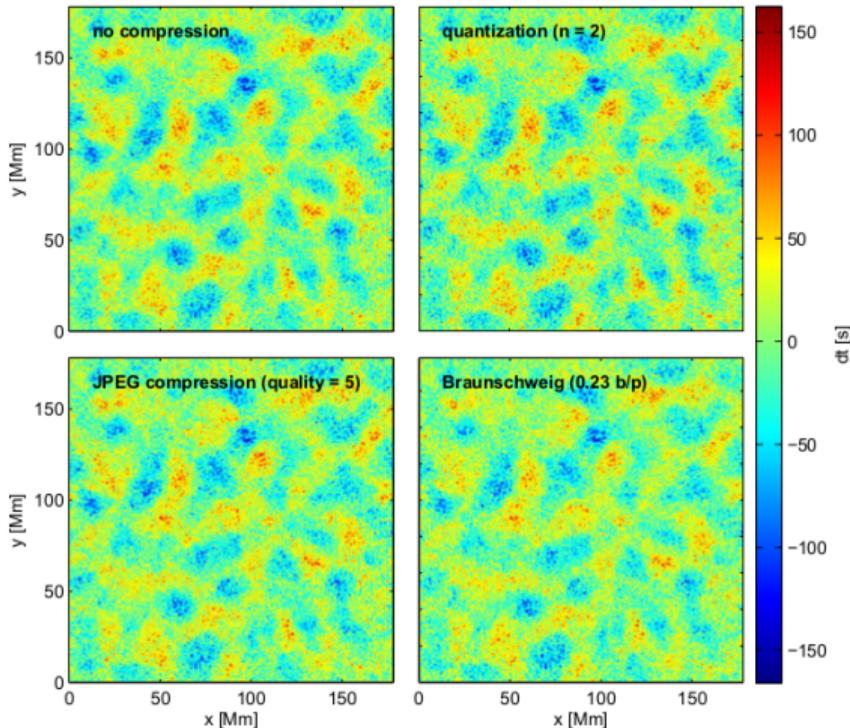


Source: v_{LOS} maps

- Compression barely affects waves
- Quantization: almost constant background noise
- JPEG and Braunschweig code:
 - Artifacts in power spectrum
 - Granulation reduced (important for LCT)

(Björn Löptien, MPS)

Travel-time Maps of Supergranulation

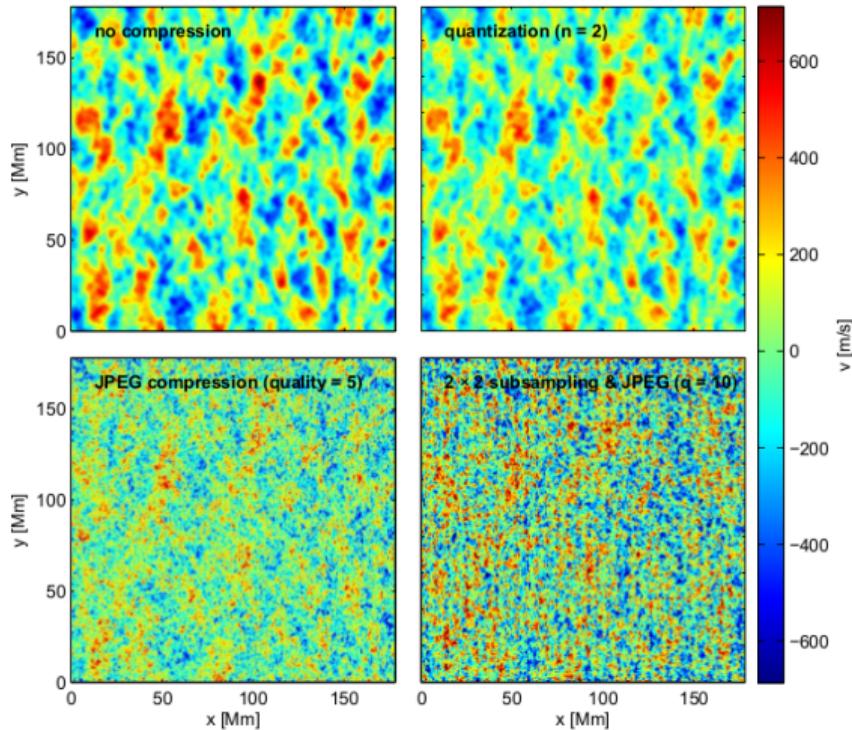


Source: v_{LOS} maps

- Divergence of supergranulation flows
- F-mode (near-surface)
Negative: outflows
Positive: inflows
- Maps from compressed data almost indistinguishable from uncompressed data No apparent correlation between noise and travel-times

(Björn Löptien, MPS)

LCT of Supergranulation



Source: I_C maps

- v_x on supergranulation scale
- Strong influence of JPEG compression and Braunschweig code

(Björn Löptien, MPS)

Summary Helioseismology (PRELIMINARY!)

- Only one simple measurement tested so far (f-mode helioseismology of supergranulation)
- Results probably not valid for helioseismology in general
- Tolerable compression efficiency depends on measurement method and science goal
- Time-distance: JPEG best method that we tested so far
- LCT: Quantization best method that we tested so far

Next steps:

- Test on near-surface latitudinal differential rotation (ongoing work)
- Test on HMI raw data (center-to-limb effects, e.g., foreshortening)